

Helical Inflation and Cosmic Strings

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OUTLINE

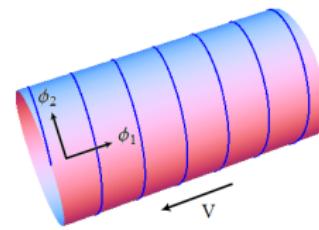
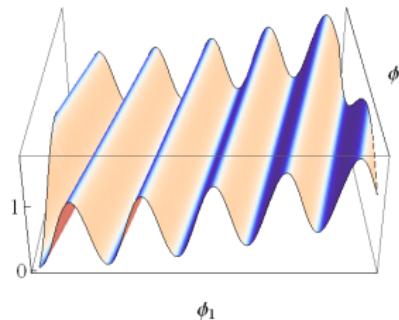
MODEL

COSMIC STRINGS

CONCLUSION

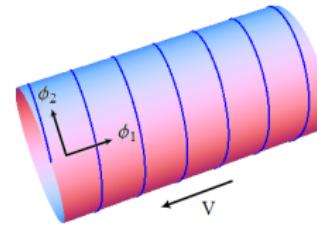
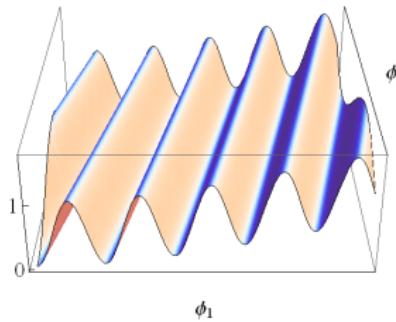
POTENTIAL

$$V(\phi_1, \phi_2) = V_0 \left\{ 1 - \cos \left(\frac{\phi_1}{f_1} \right) + A \left[1 - \cos \left(\frac{\phi_1}{f'_1} - \frac{\phi_2}{f_2} \right) \right] \right\}$$



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K. Choi, H. Kim and S. Yun, arXiv:1404.6209 [hep-th]

T. Higaki and F. Takahashi, arXiv:1404.6923 [hep-th]

S. -H. H. Tye and S. S. C. Wong, arXiv:1404.6988 [astro-ph.CO]

R. Kappl, S. Krippendorf and H. P. Nilles, arXiv:1404.7127 [hep-th]

I. Ben-Dayan, F. G. Pedro and A. Westphal, arXiv:1404.7773 [hep-th]

C. Long, L. McAllister and P. McGuirk, arXiv:1404.7852 [hep-th]

MOTIVATION

$$V_{eff}(\phi) \simeq V_0 \left[1 - \cos \left(\frac{\phi}{f_{eff}} \right) \right], \quad f_{eff} \simeq \frac{f_1 f_2}{f'_1}$$

- ▶ Large field inflation
- ▶ Sub-Planckian $f_i \rightarrow$ super-Planckian f_{eff}

J. E. Kim, H. P. Nilles and M. Peloso, JCAP **0501** (2005) 005 [hep-ph/0409138]

- ▶ Free from UV corrections due to axion symmetry
 $\phi \rightarrow \phi + 2\pi f$
- ▶ Realization in SUGRA

R. Kallosh, A. Linde and B. Vercnocke, arXiv:1404.6244 [hep-th]

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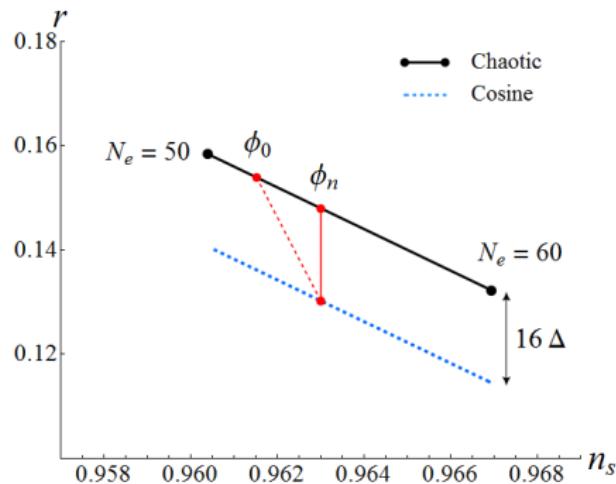
R. Kallosh, A. Linde and B. Vercnocke, arXiv:1404.6244 [hep-th]

$N_e = 60 - 50$	n_s	r	$\hat{\Delta} = r + 4(n_s - 1)$
Chaotic Inflation	0.967 – 0.960	0.132 – 0.158	0
Helical Model ($\frac{f_1 f_2}{f'_1} = 20M_{pl}$)	0.966 – 0.960	0.124 – 0.150	-0.011
Helical Model ($\frac{f_1 f_2}{f'_1} = 14M_{pl}$)	0.966 – 0.960	0.115 – 0.141	-0.020
Experimental Data	0.9607 ± 0.0063	$0.16^{+0.06}_{-0.05}$	

COSINE FEATURE

$$\hat{\Delta} = 16\Delta = r + 4(n_s - 1)$$

$$\Delta_{\phi^2} = 0, \quad \Delta_{cosine} = -\frac{1}{4f^2}$$



COSINE FEATURE

$V(\phi)$	$\frac{1}{2}m^2\phi^2$	$V_0 \left[1 - \cos \left(\frac{\phi}{f}\right)\right]$
$\epsilon = \frac{1}{2} \left(\frac{V'}{V} \right)^2$	$\frac{1}{16}r$	$\frac{1}{16}r$
$\eta = \left(\frac{V''}{V} \right)^2$	$\frac{1}{16}r$	$\frac{1}{16}r + 2\Delta$
$\xi^2 = \frac{V'V'''}{V^2}$	0	$\frac{1}{2}r\Delta$
$\omega^3 = \frac{V'^2V''''}{V^3}$	0	$\frac{1}{32}r^2\Delta + r\Delta^2$
$n_s - 1 = 2\eta - 6\epsilon$	$-\frac{1}{4}r$	$-\frac{1}{4}r + 4\Delta$
$n_t = -2\epsilon$	$-\frac{1}{8}r$	$-\frac{1}{8}r$
$\frac{dn_s}{d\ln k} = -16\epsilon\eta + 24\epsilon^2 + 2\xi^2$	$\frac{1}{32}r^2$	$\frac{1}{32}r^2 - r\Delta$
$\frac{dn_t}{d\ln k} = -4\epsilon\eta + 8\epsilon^2$	$\frac{1}{64}r^2$	$\frac{1}{64}r^2 - \frac{1}{2}r\Delta$
$\frac{d^2n_s}{d\ln k^2} = -192\epsilon^3 + 192\epsilon^2\eta - 32\epsilon\eta^2 - 24\epsilon\xi^2 + 2\eta\xi^2 + 2\omega^3$	$-\frac{1}{128}r^3$	$-\frac{1}{128}r^3 + \frac{3}{8}r^2\Delta - 4r\Delta^2$

COSINE FEATURE

Independent ways of measuring Δ ,

$$8\Delta^2 = \frac{1}{2}(n_s - 1)^2 - \frac{dn_s}{d \ln k}, \quad \hat{\Delta} = 16\Delta = \frac{r}{2} - \frac{32}{r} \frac{dn_t}{d \ln k}$$

COSINE FEATURE

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$$8\Delta^2 = \frac{1}{2}(n_s - 1)^2 - \frac{dn_s}{d \ln k}, \quad \hat{\Delta} = 16\Delta = \frac{r}{2} - \frac{32}{r} \frac{dn_t}{d \ln k}$$

Bounds on $\hat{\Delta}$,

$$m^2 = \frac{V_0}{f^2} \simeq 4 \times 10^{-11} M_{pl}^2$$

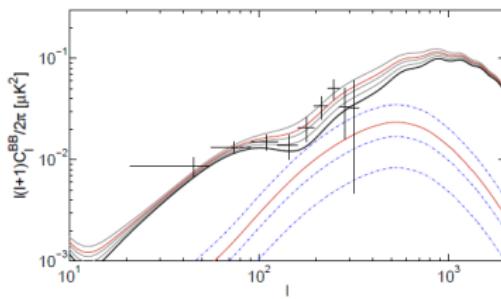
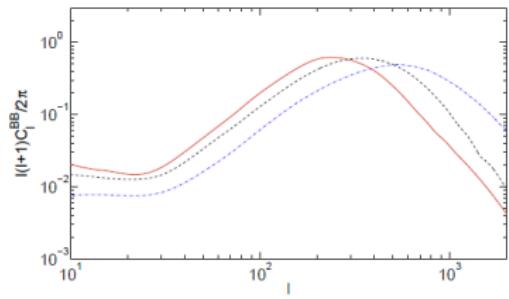
$$\hat{\Delta} = -\frac{4M_{pl}^2}{f^2} = -0.03 \frac{V_{infl}}{V_0} = -0.03 \frac{(2 \times 10^{16} GeV)^4}{V_0}$$

$$M_{pl} \gg M_S \geq M_{GUT} \simeq 2 \times 10^{16} GeV$$

COSMIC STRINGS

- ▶ $V_{infl} \simeq M_{GUT}$
- ▶ Many axions and $U(1)$ gauge fields \rightarrow vortices
- ▶ Evolving to scaling network, properties reliably calculated

COSMIC STRINGS



J. Lizarraga, J. Urrestilla, D. Daverio, M. Hindmarsh, M. Kunz and A. R. Liddle, Phys. Rev. Lett. **112** (2014) 171301 [arXiv:1403.4924 [astro-ph.CO]]

A. Moss and L. Pogosian, Phys. Rev. Lett. **112** (2014) 171302 [arXiv:1403.6105 [astro-ph.CO]]

CONCLUSION

- ▶ Natural inflation looks natural
- ▶ Looking forward to see more experimental results
- ▶ Cosmic strings
- ▶ GUT again?