

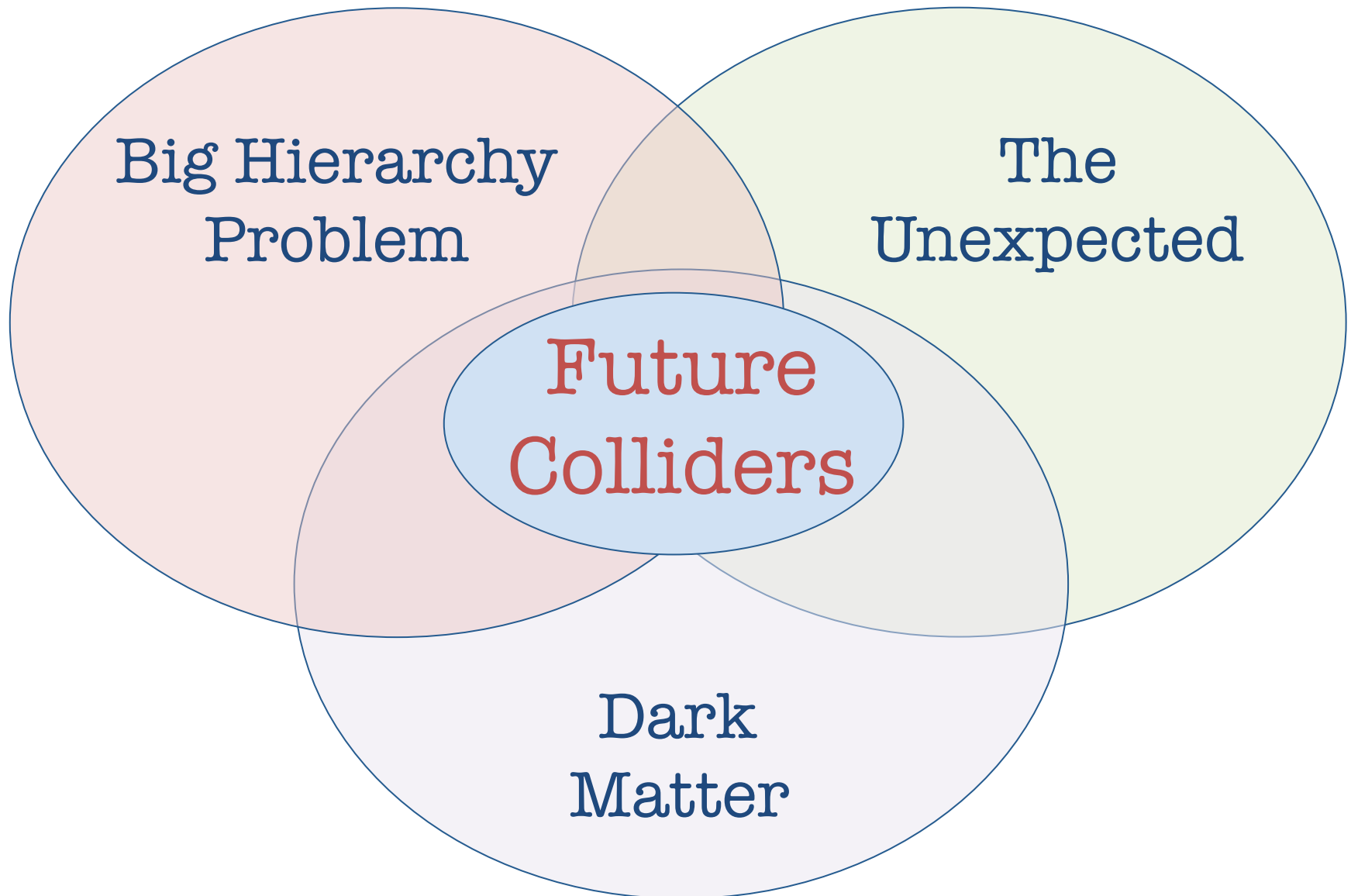
New Physics at 100 TeV

HKUST Jockey Club IAS
Jan 19th 2015

Matthew McCullough



Beyond the Standard Model



Dark Matter at 100 TeV

Despite overwhelming evidence for its existence, the particle nature of dark matter is unknown.

Cosmology provides a strong motivation for direct and collider searches...

- Thermal freeze-out predicts observed abundance for:

$$M_{DM} \sim \mathcal{O}(\text{few GeV}) \rightarrow \mathcal{O}(10\text{'s TeV})$$

↑
Cosmological constraints

↑
Unitarity bounds

Motivates dark matter searches in ballpark of 100 TeV collider independent of hierarchy problem.

Simplified Dark Matter Models

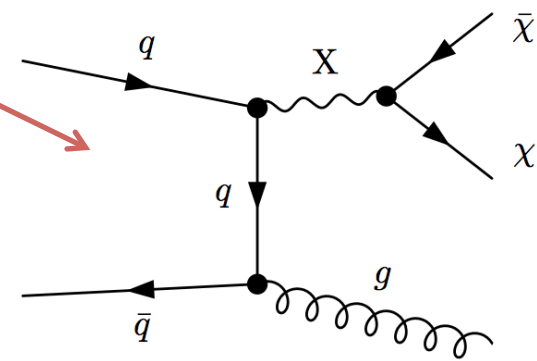
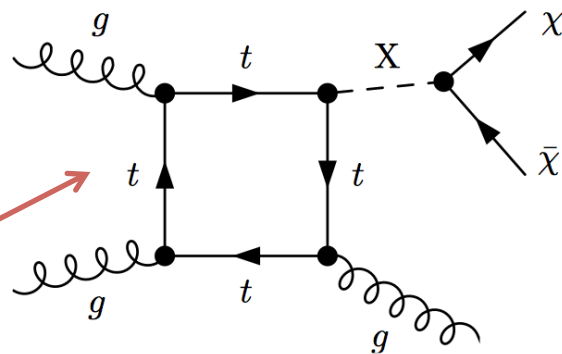
Write down simple scenarios to model production of dark matter at colliders:

$$\mathcal{L}_S \supset - \sum_q c_S \lambda_{h,q} S \bar{q} q - \frac{1}{2} m_{\text{MED}}^2 S^2 + \mathcal{L}(S, \bar{\chi}, \chi),$$

$$\mathcal{L}_P \supset - \sum_q i c_P \lambda_{h,q} P \bar{q} \gamma^5 q - \frac{1}{2} m_{\text{MED}}^2 P^2 + \mathcal{L}(P, \bar{\chi}, \chi),$$

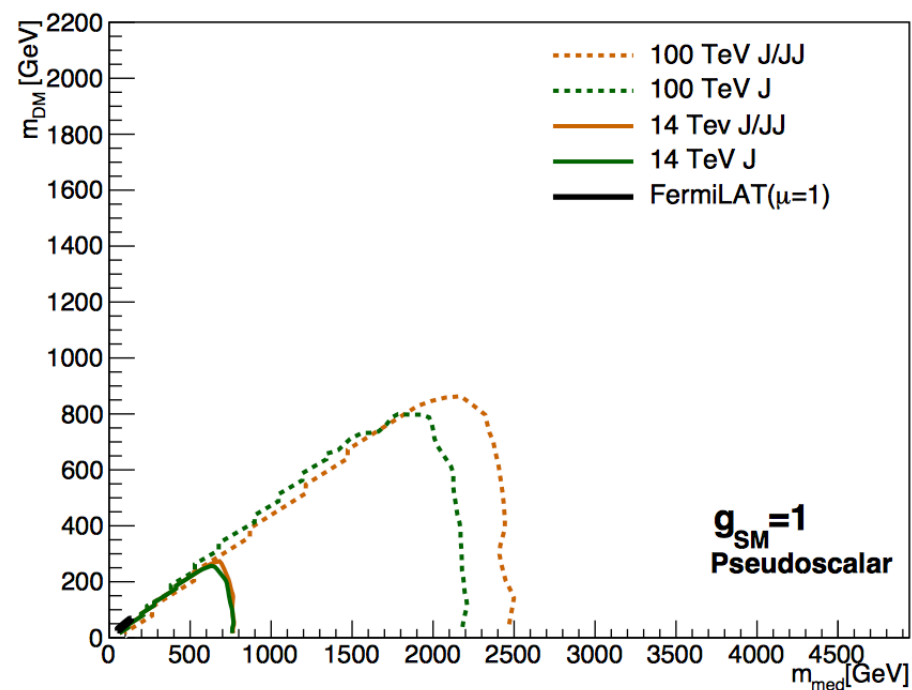
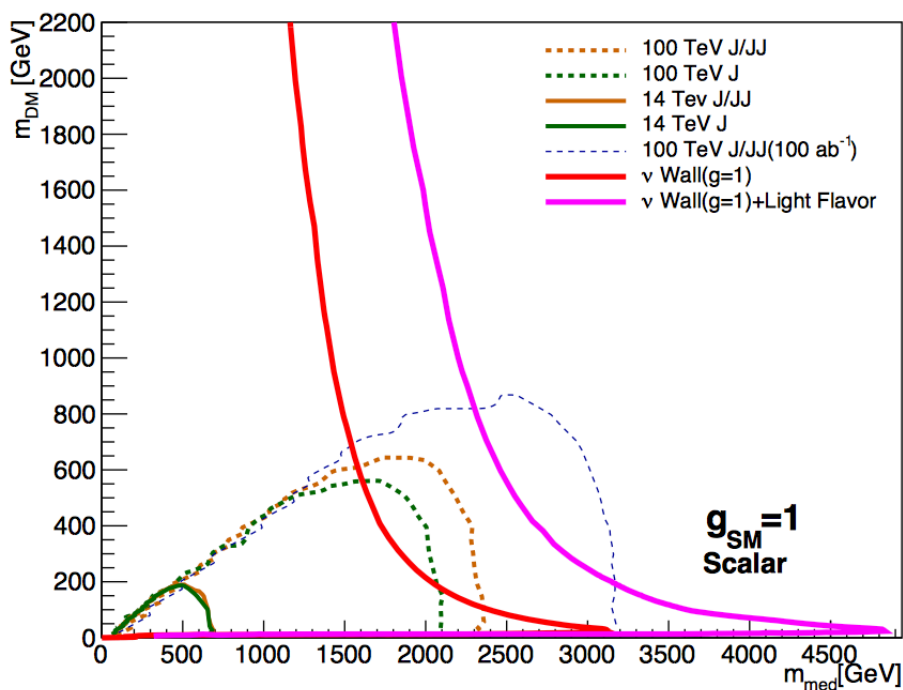
$$\mathcal{L}_V \supset - \sum_q c_V V_\mu \bar{q} \gamma^\mu q - \frac{1}{2} m_{\text{MED}}^2 V_\mu V^\mu + \mathcal{L}(V, \bar{\chi}, \chi),$$

$$\mathcal{L}_A \supset - \sum_q c_A A_\mu \bar{q} \gamma^\mu \gamma^5 q - \frac{1}{2} m_{\text{MED}}^2 A_\mu A^\mu + \mathcal{L}(A, \bar{\chi}, \chi),$$



Simplified Dark Matter Models

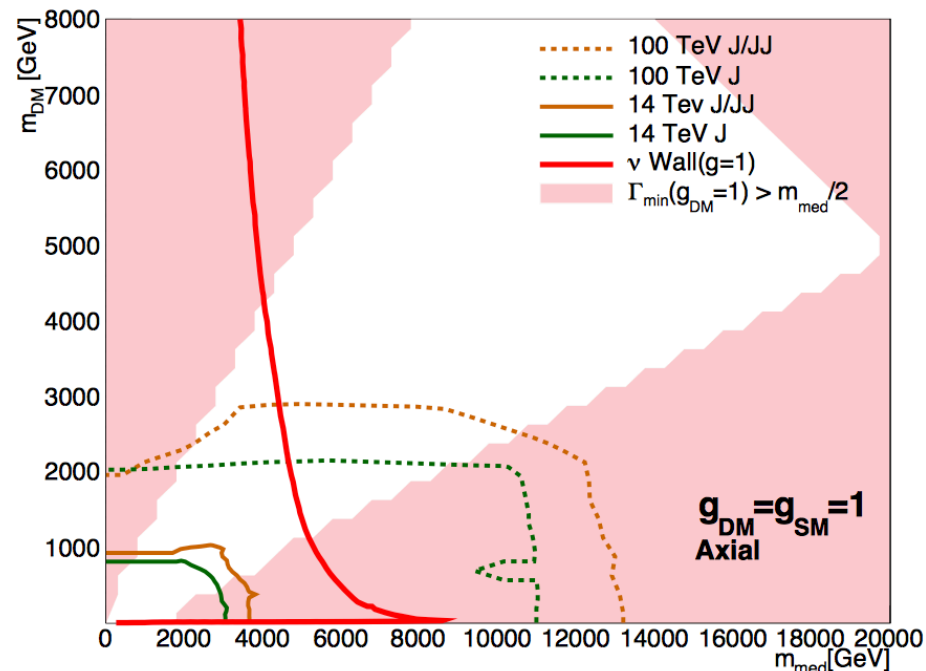
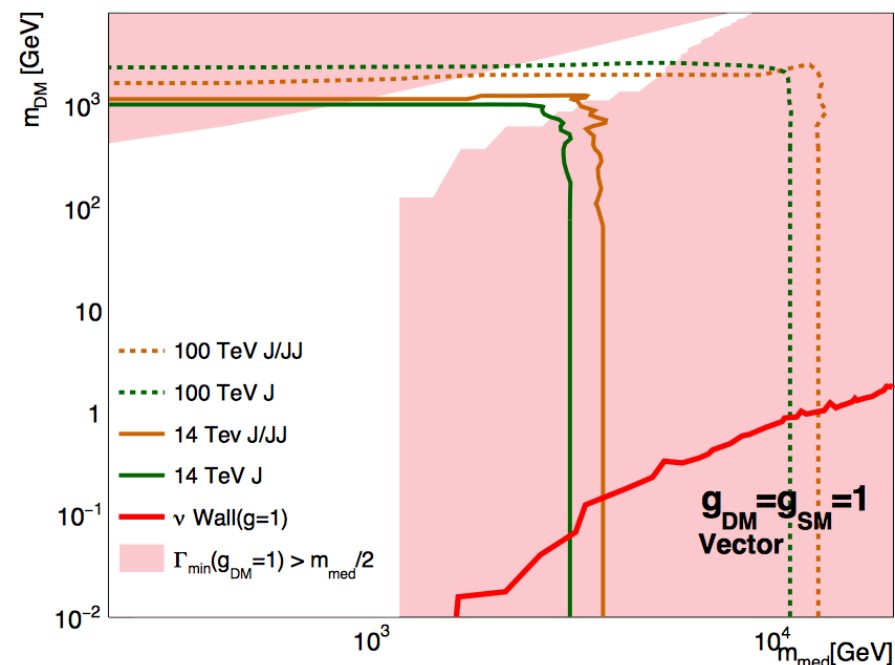
Coverage of simplified model parameter space is extended at 100 TeV collider for all mediators:



100 TeV Study: Harris, Khoze, Spannowsky, Williams, 2015.

Simplified Dark Matter Models

Coverage of simplified model parameter space is extended at 100 TeV collider for all mediators:



Littlest Simplified Model?

The Higgs Portal:

ϕ is the dark matter?

$$\mathcal{L} = \mathcal{L}_{SM} - \frac{1}{2}\partial_\mu\phi\partial^\mu\phi - \frac{1}{2}M^2\phi^2 - c_\phi|H|^2\phi^2$$

The Higgs itself could be the “mediator” to the dark sector!

Higgs Portal at 100 TeV

If the singlet is DM, i.e. only pair-produced, we have:

$$\mathcal{L} = \mathcal{L}_{SM} - \frac{1}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{2} M^2 \phi^2 - c_\phi |H|^2 \phi^2$$

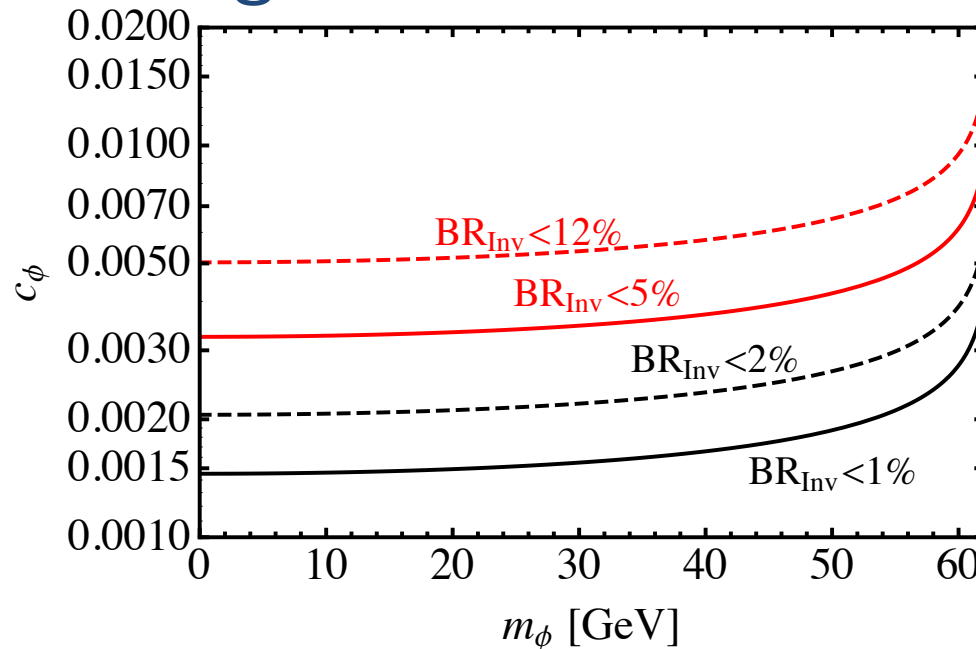
Usual limits on-shell: $h \rightarrow \phi\phi \Rightarrow \Gamma_{h \rightarrow \text{inv}} \neq 0$

Relevant if we have: $m_\phi < m_h/2 < 62 \text{ GeV}$

Can be constrained by global coupling fit, however MET+(di-)jet signatures also promising for direct search.

Higgs Portal at 100 TeV

Coupling sensitivity possible with variety of invisible branching limits.

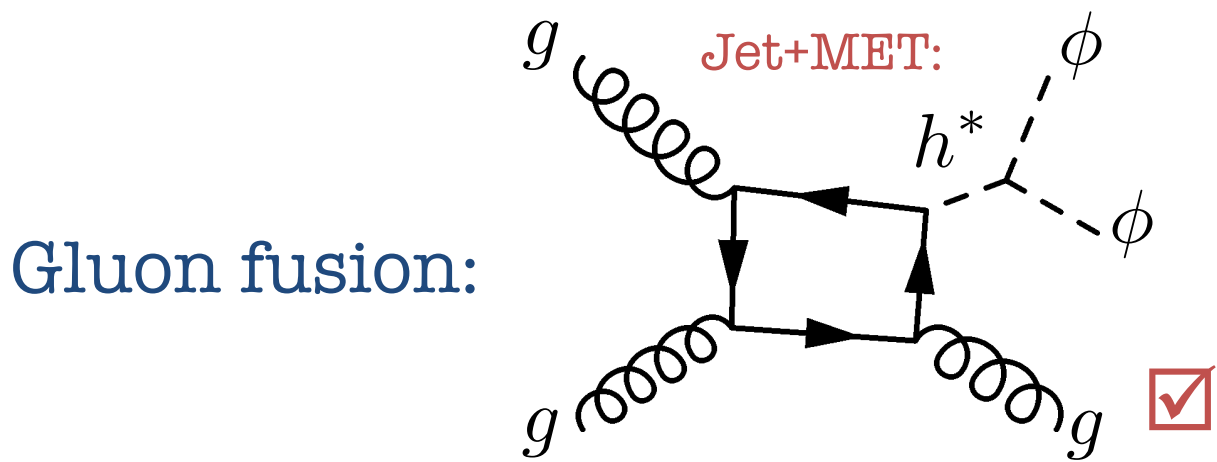
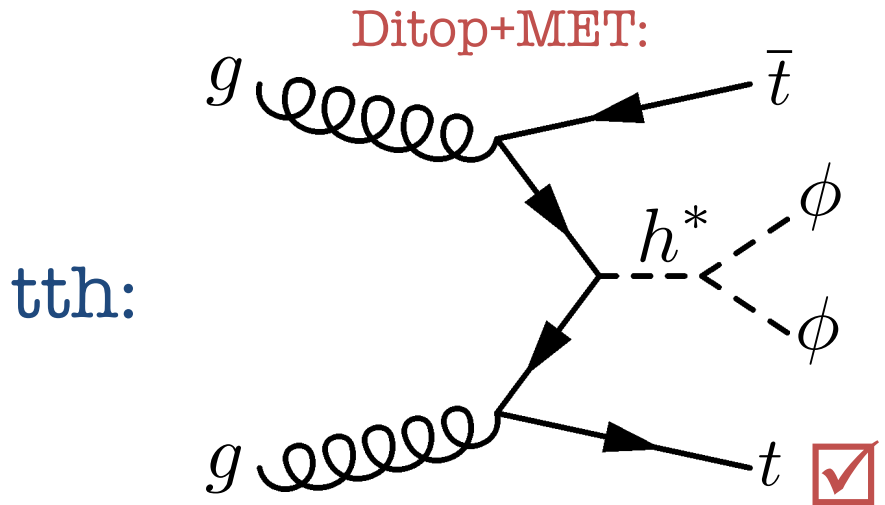
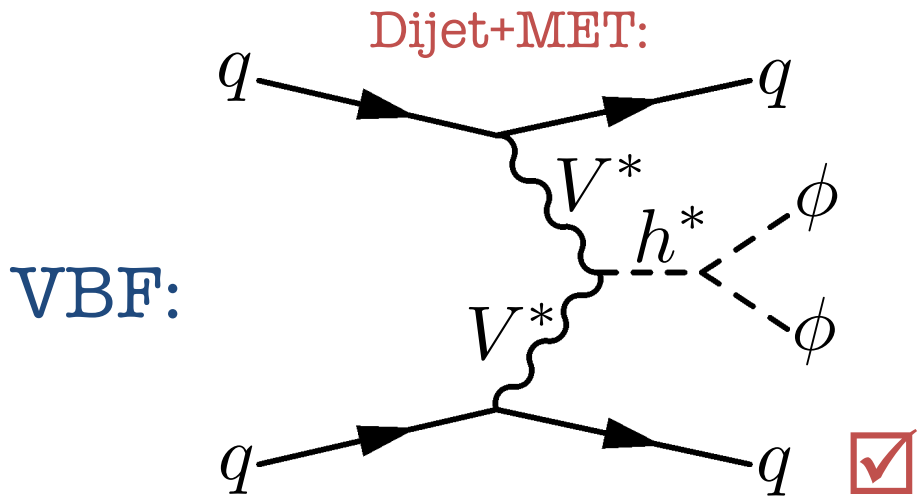


100 TeV capability depends on coupling precision.

What about $m_\phi > m_h/2$? This is the majority of parameter space. What can we do?

Higgs Portal

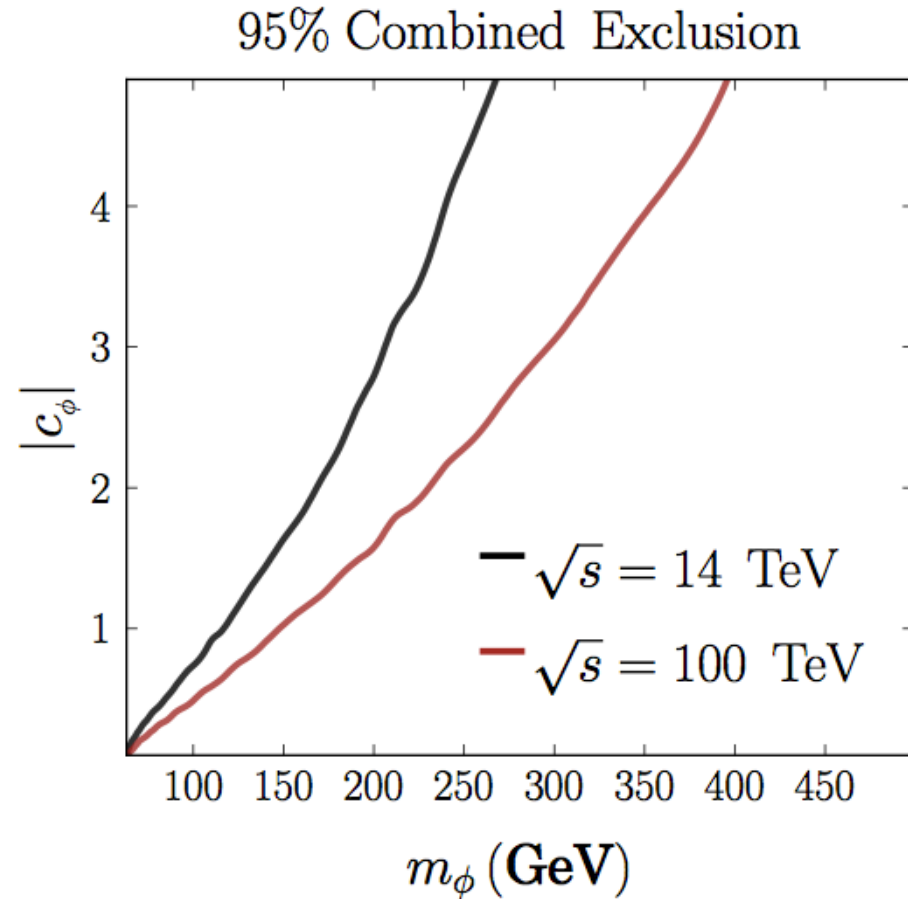
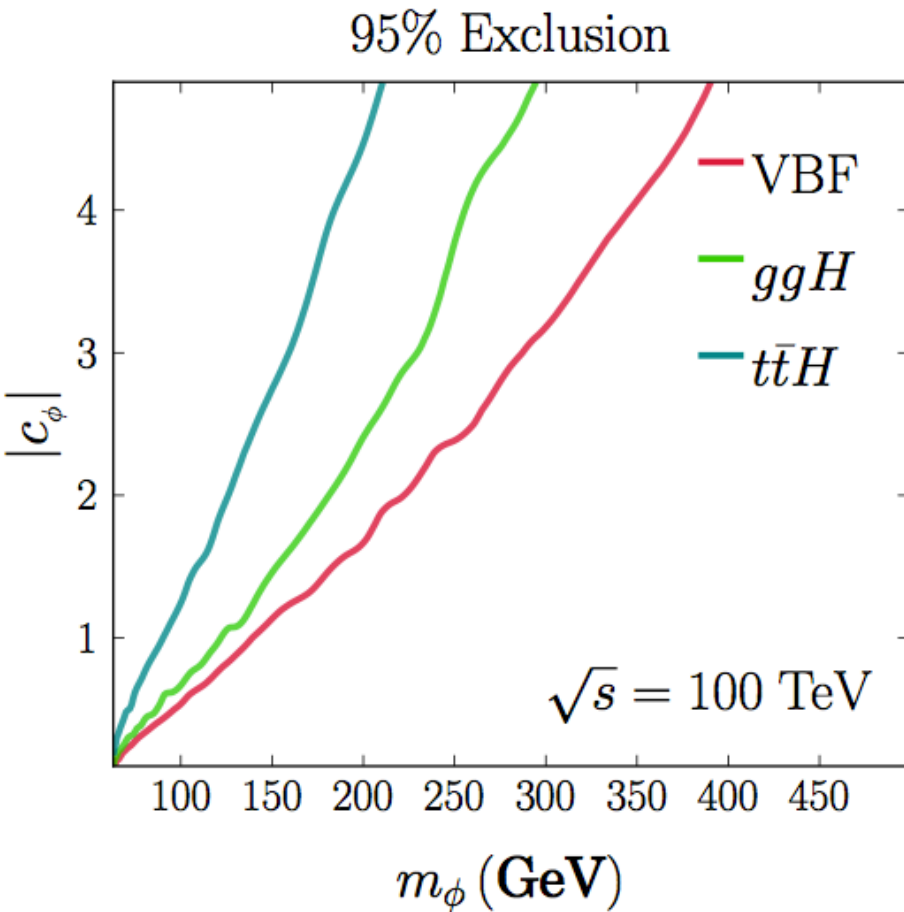
Want more coverage! For $m_\phi > m_h/2$ must go off-shell...



Note: Higgs EFT overestimates rate so full loop necessary.

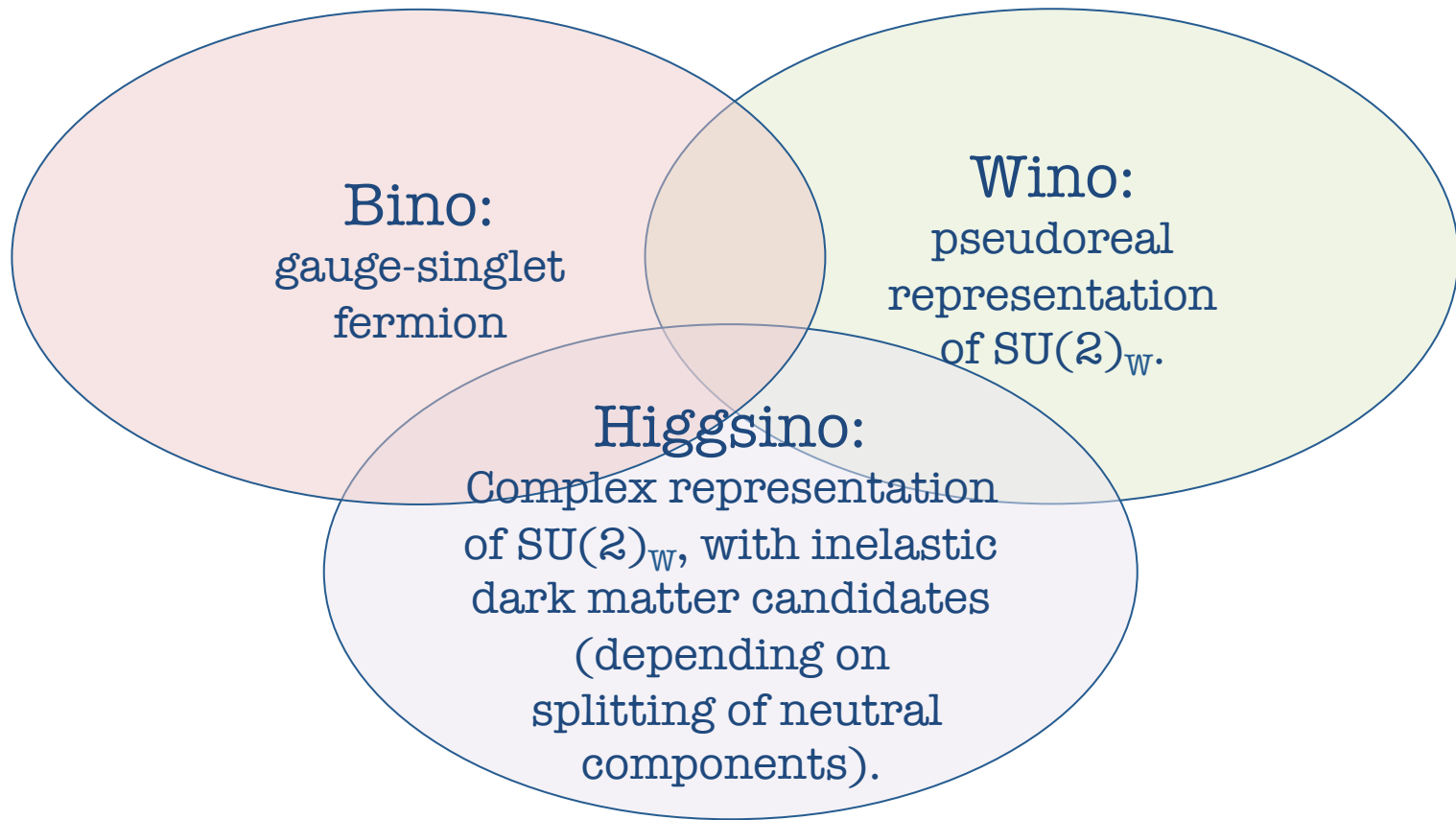
Higgs Portal

Different channel and different colliders:



Neutralino Dark Matter

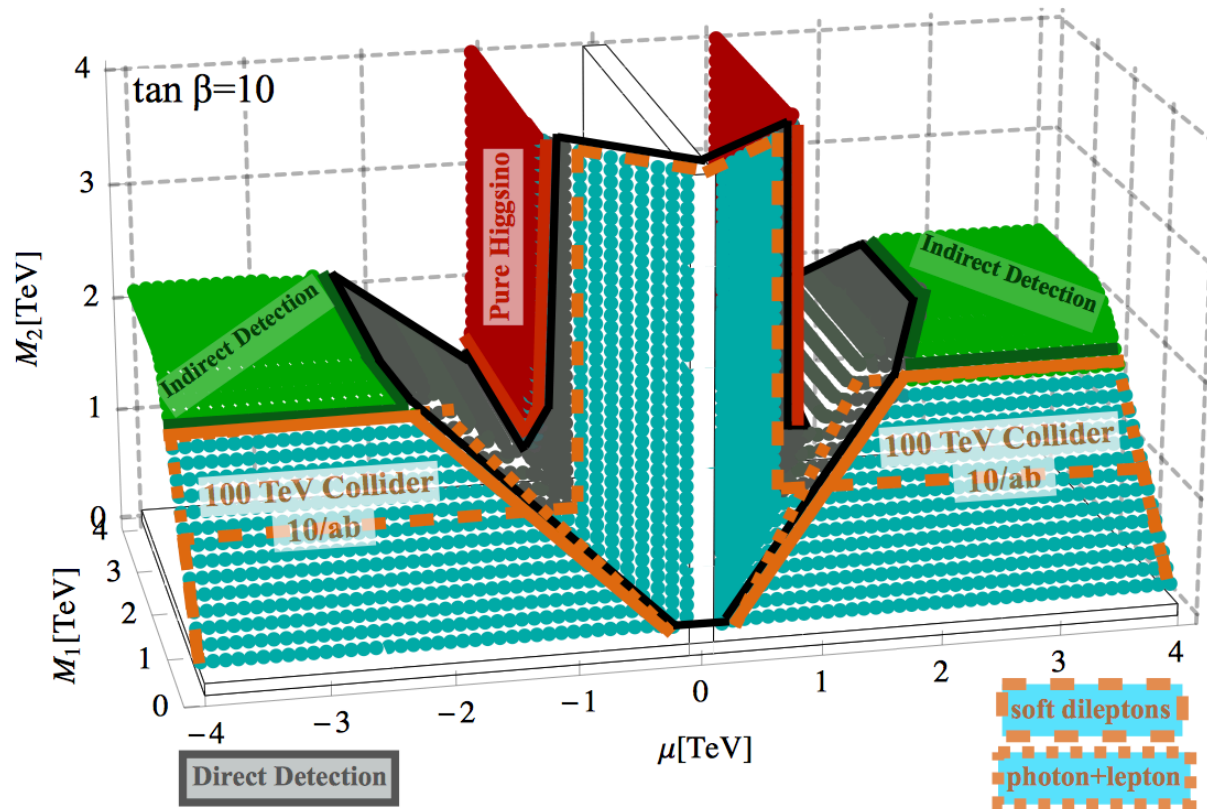
Neutralino dark matter is a compelling ingredient of the SUSY setup, including Mini-Split SUSY.



Direct and indirect detection interplay non-trivial.

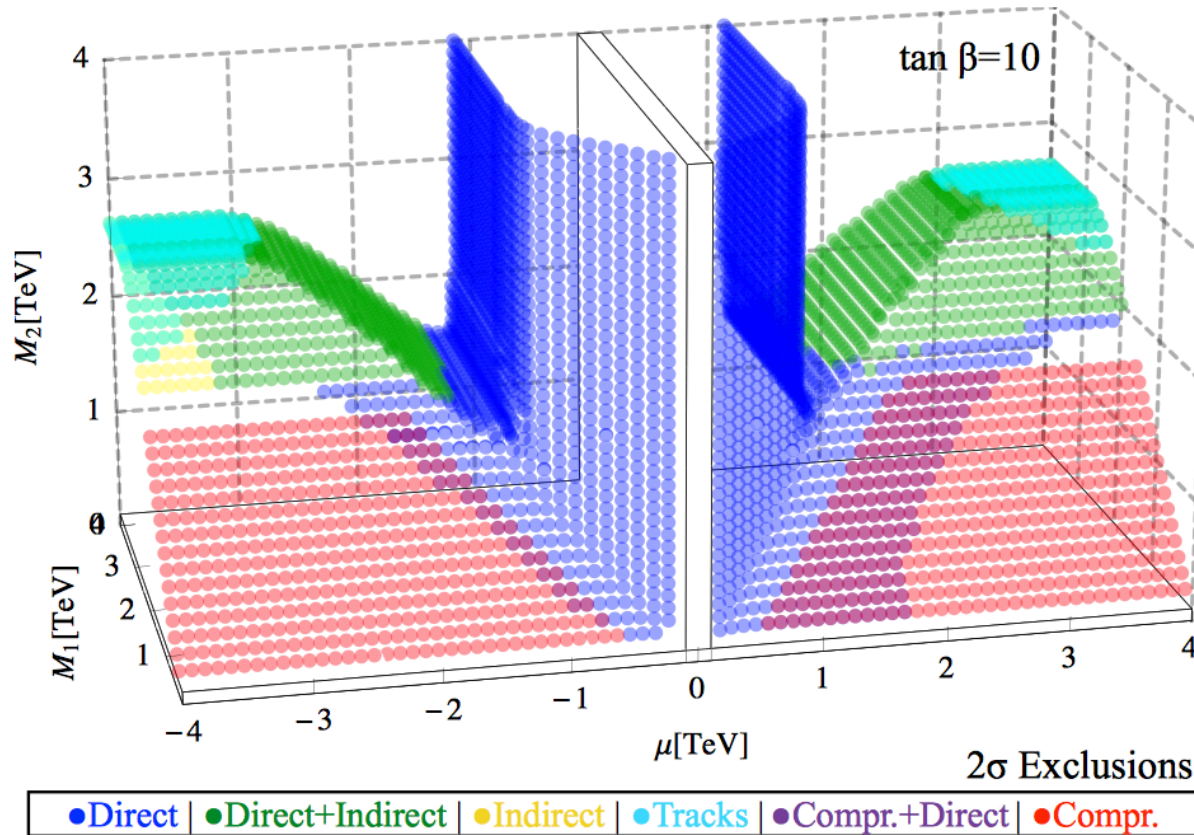
Relic Neutralino Surface

Considering just the electroweakino sector of MSSM. Collider signatures considered: MET + Jet and either soft dileptons or lepton+photon



Relic Neutralino Surface

Considering just the electroweakino sector of MSSM. Collider signatures considered: Disappearing charged tracks and previous slide.



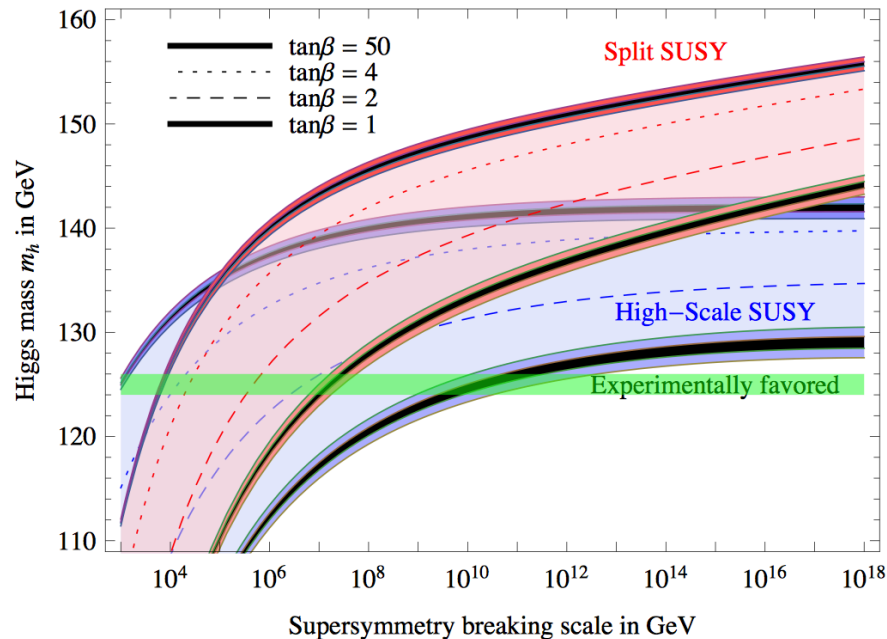
Relic Neutralino Surface

See also: Gori, Jung, Wang, Wells. 2014.
And: Acharya, Bozek, Pongkitivanichkul, Sakurai. 2014.
And: Cirelli, Sala, Taoso. 2014.
And: Buchmuller, Citron, Ellis, Guha, Marrouche, Olive, de
Vries, Zheng. 2015.

Supersymmetry

SUSY predicts collider signatures both as a potential solution to the hierarchy problem and also if the weak scale is tuned (Mini-Split).

Predicted range for the Higgs mass



Degrassi, Di Vita,
Elias-Miro, Espinosa,
Giudice, Isidori, Strumia

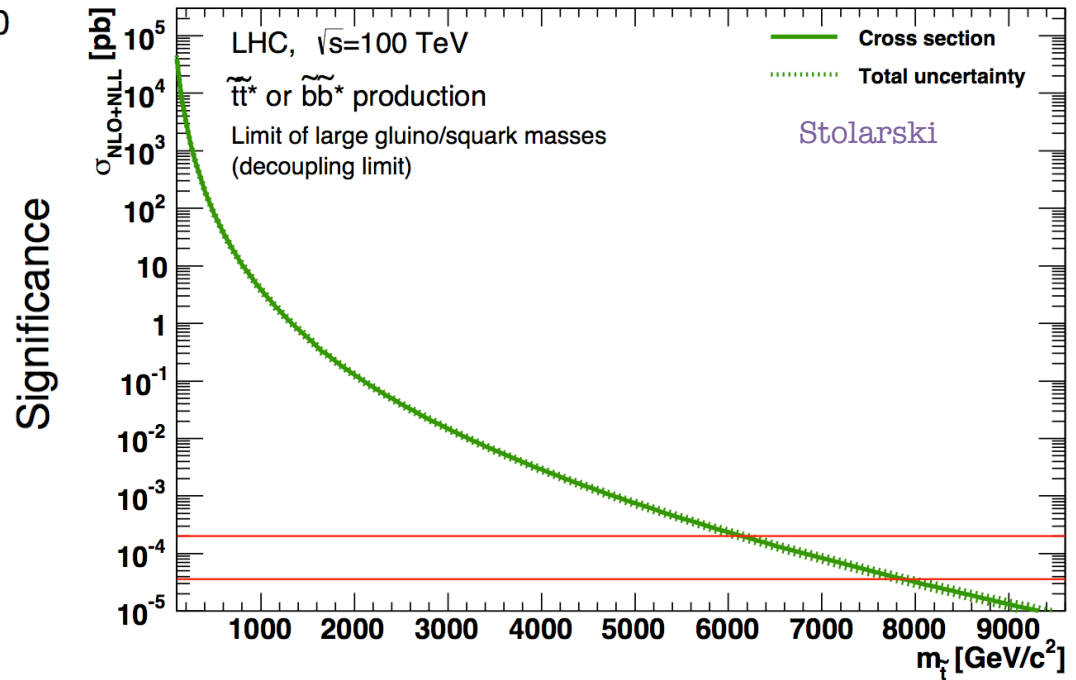
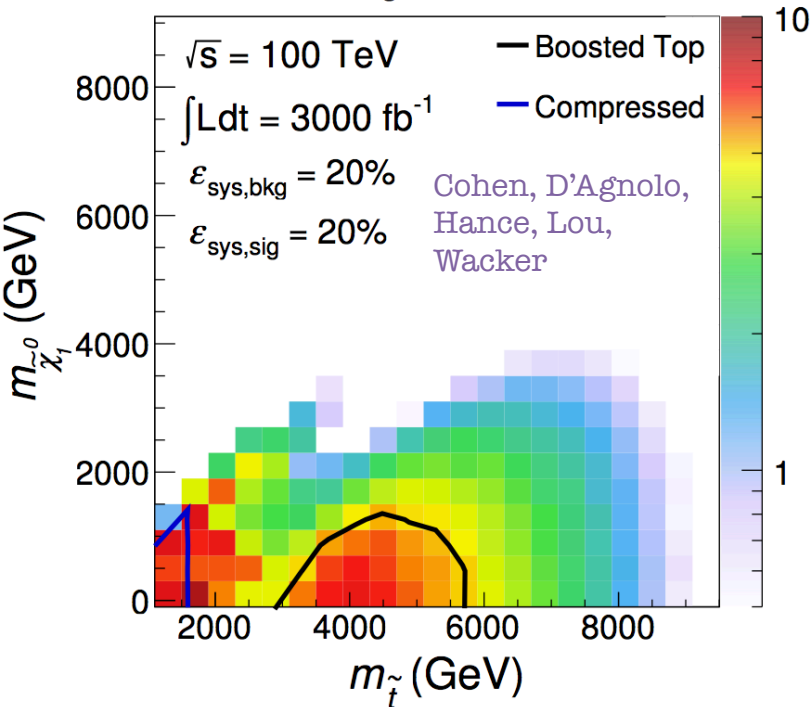
Higgs mass in simplest scenario (MSSM) already provides a ballpark expectation for squarks.

Supersymmetry

Stop squarks are most relevant for weak scale naturalness, thus direct searches directly impact

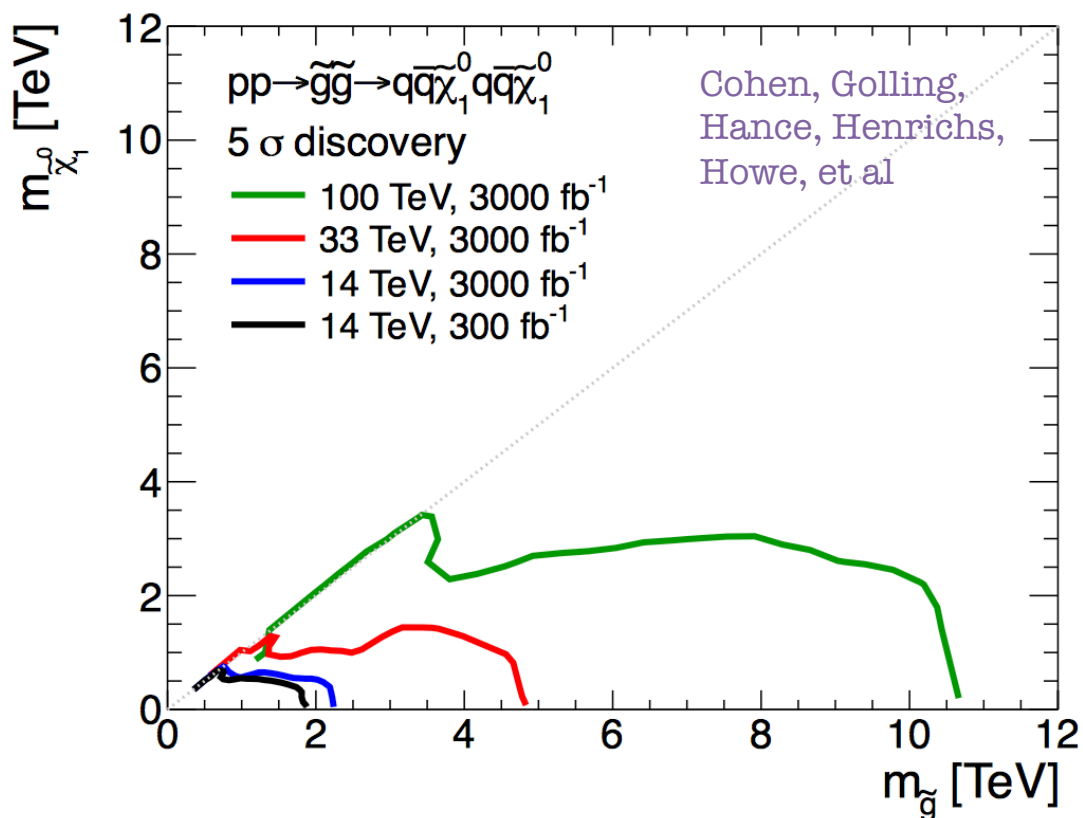
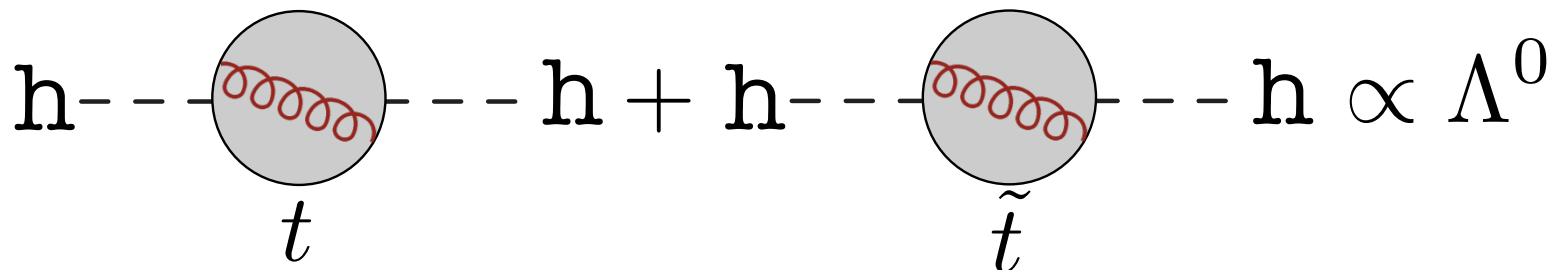
tuning: $h \dashv\dashv \overset{\circ}{\propto \Lambda^2} \dashv\dashv h + h \dashv\dashv \overset{\circ}{\propto \Lambda^2} \dashv\dashv h \propto \Lambda^0$
 t \tilde{t}

CL_s Discovery



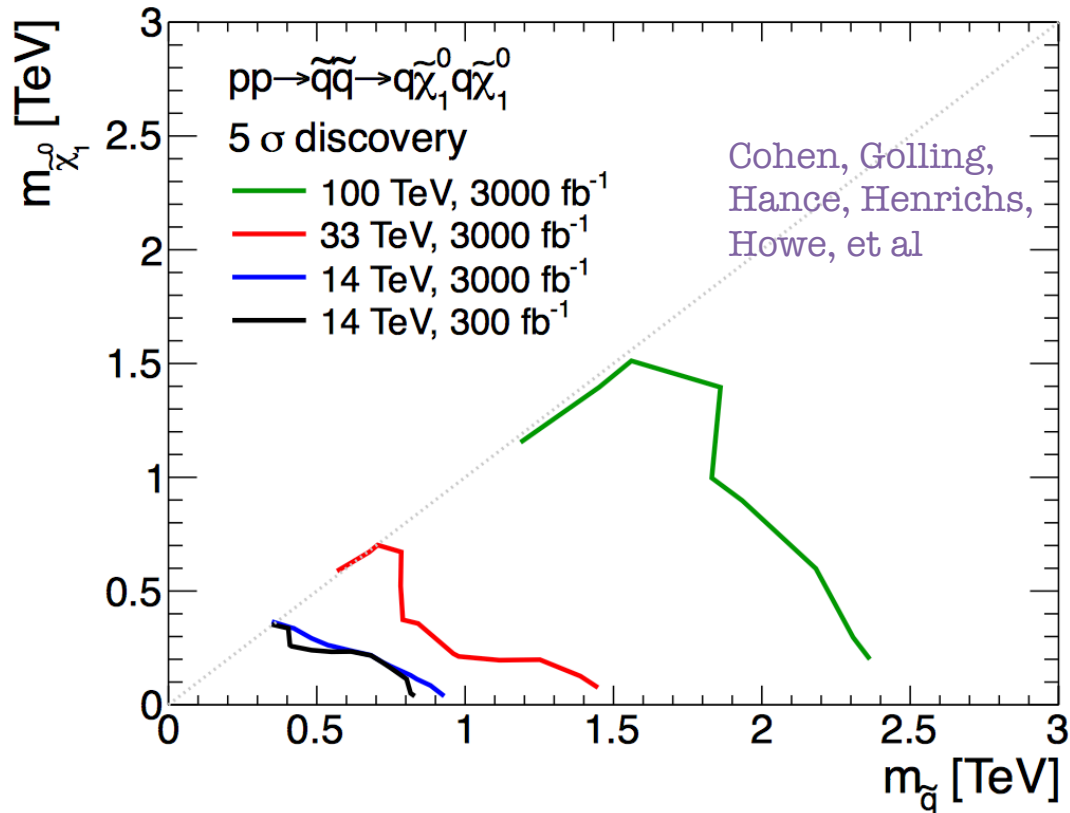
Supersymmetry

Next in line at two loops are the gluinos



Supersymmetry

First two generation squarks also discoverable

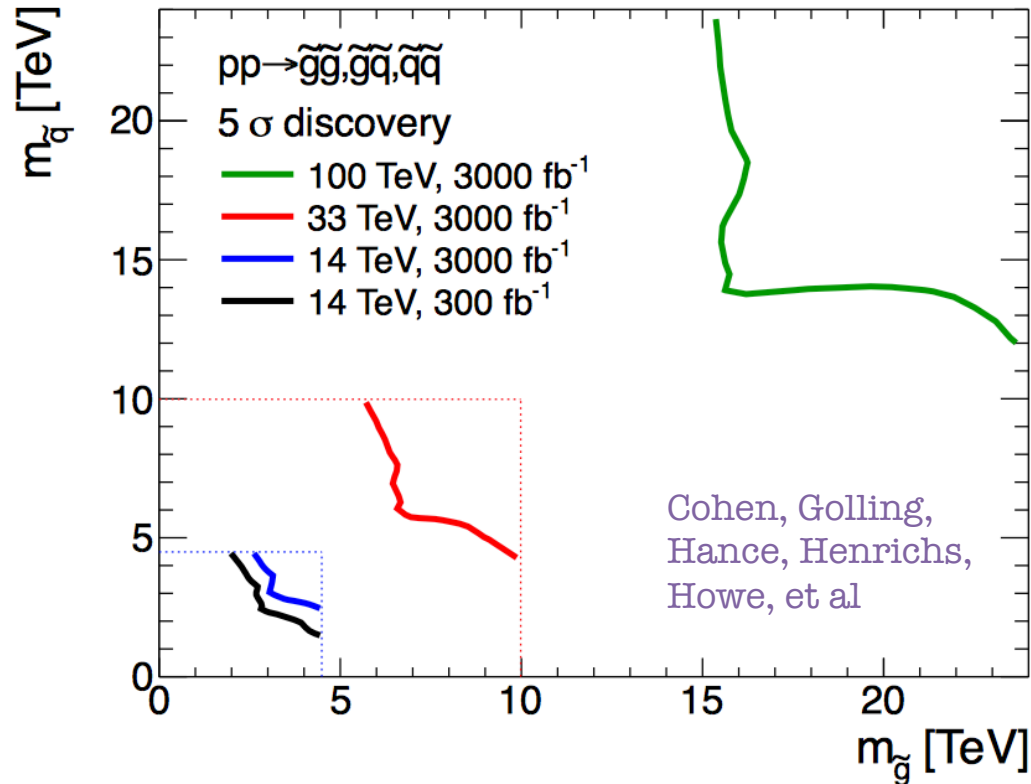


Reach suppressed compared to gluinos as decoupling gluinos suppresses cross section.

Reach increases when gluinos reintroduced...

Supersymmetry

Combining gluino and squark limits onto one plane:

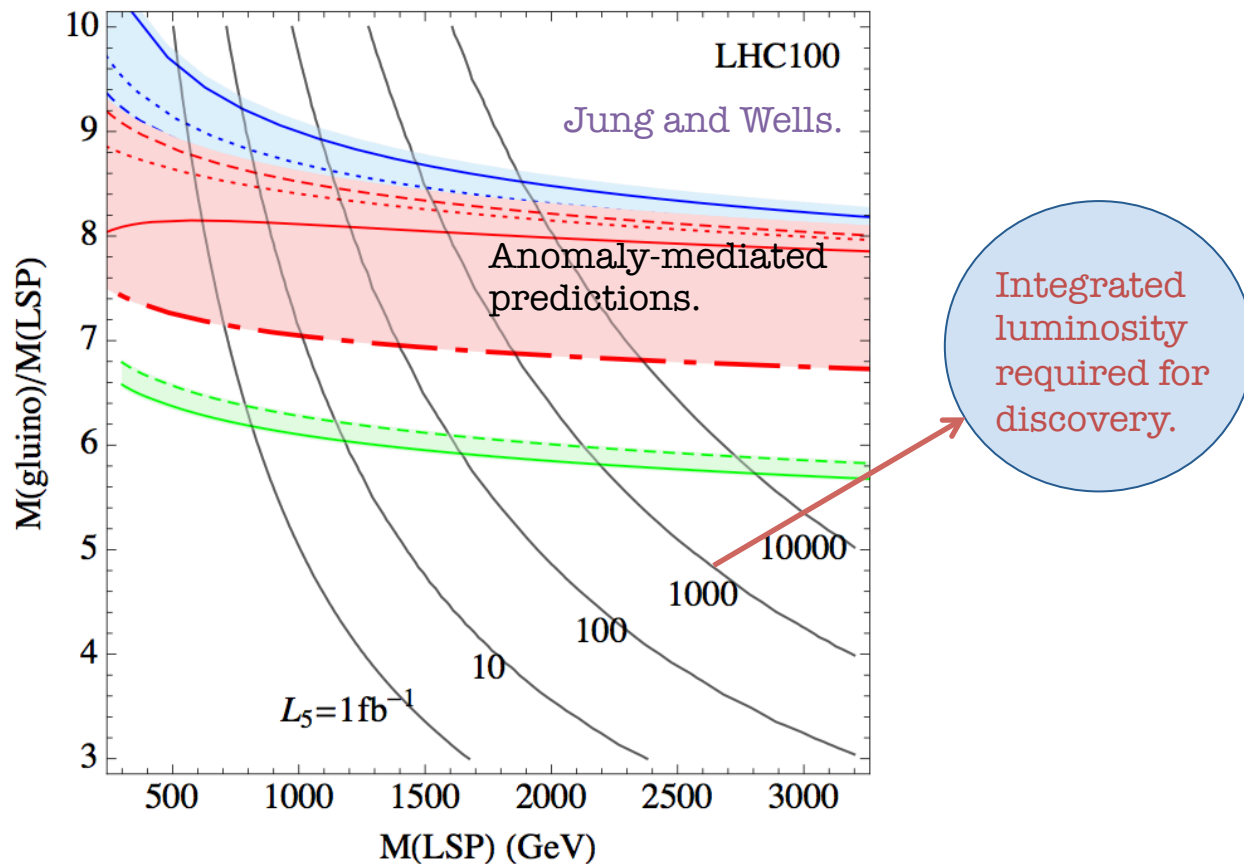


Reach in combination stronger than individual reach to due enhanced production rates with all states present.

Supersymmetry

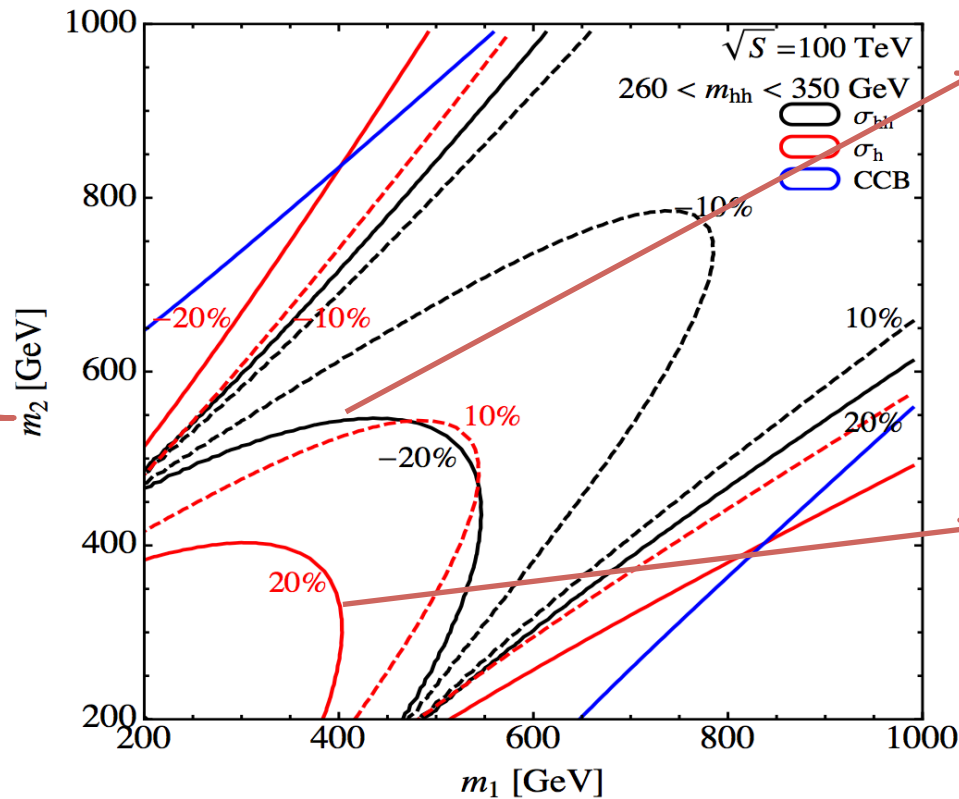
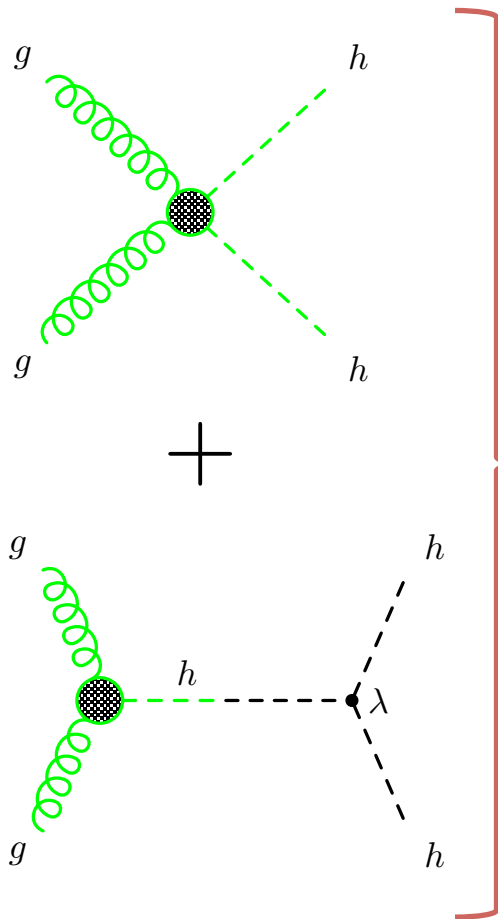
Mini-split SUSY: scalars decoupled and only fermions remains light due to R-symmetry.

Motivations: Unification, Dark Matter



Supersymmetry

Can also indirectly probe presence of SUSY states from precision Higgs measurements.



Single Higgs measurements probe Higgs-gluon coupling.

Higgs pair measurements probe further structure of Higgs-gluon coupling.

Composite Higgs

The weak scale may also be natural if Higgs is a composite of strongly interacting states.

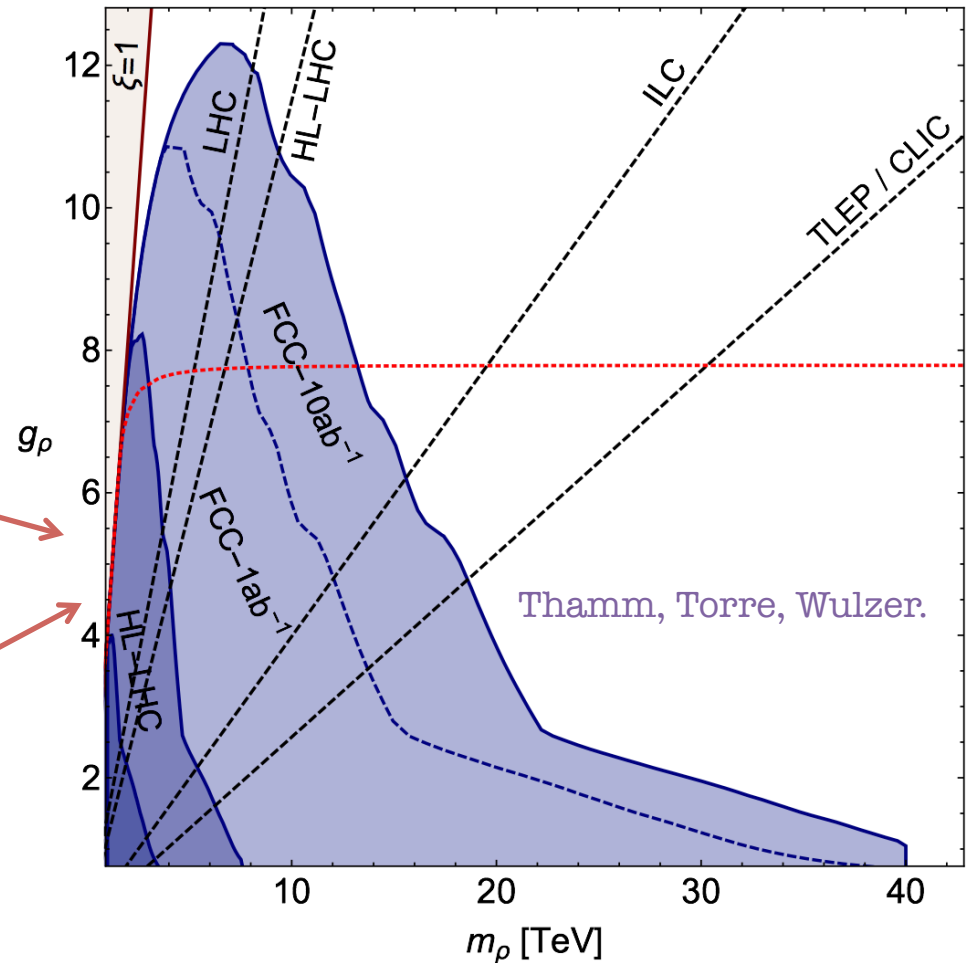
Signatures include:

- Modified Higgs couplings
- Modifications to electroweak sector, influencing precision electroweak
- New heavy resonances coupled to the SM states

$$\rho \rightarrow l\bar{l}$$

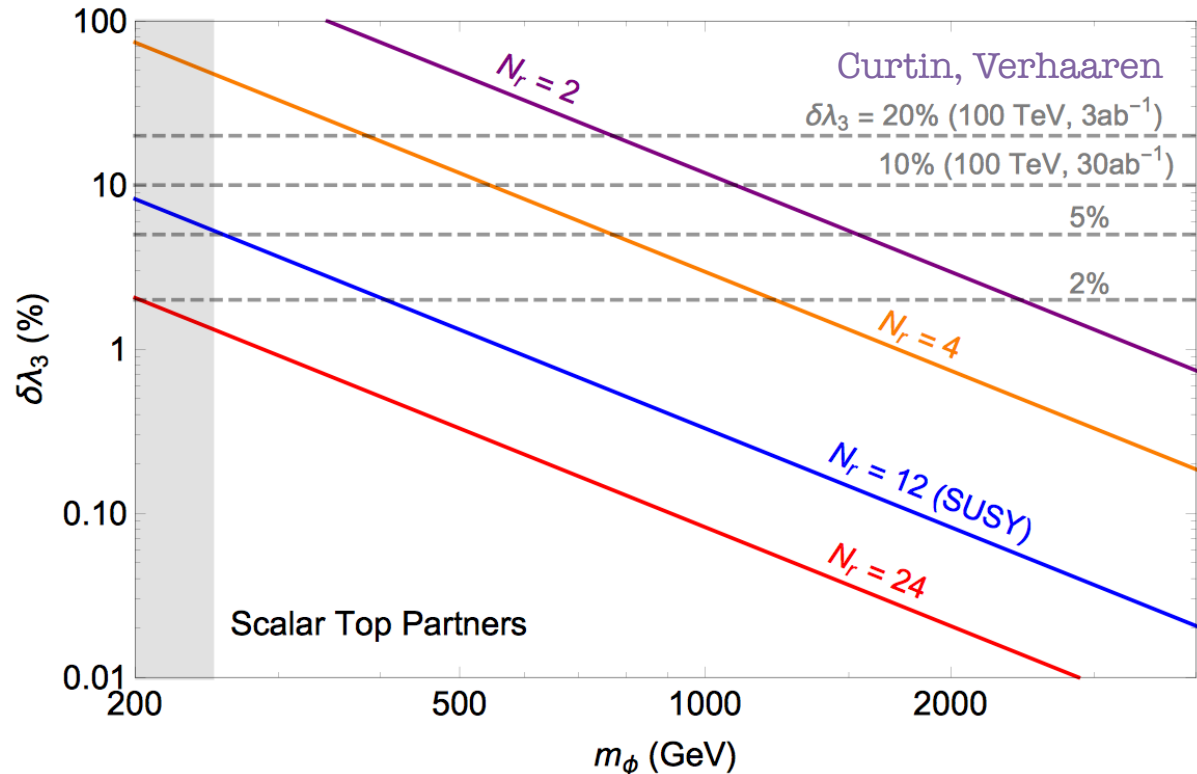
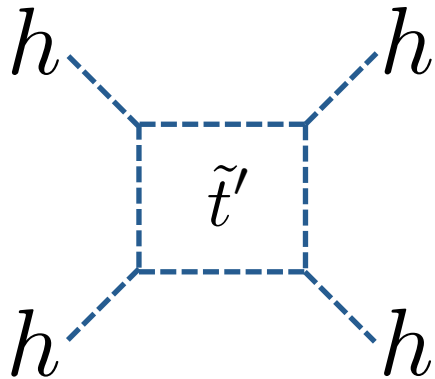
and

$$\rho \rightarrow WZ \rightarrow \nu + 3l$$



Neutral Naturalness

There are exotic theories, such as Twin Higgs or Folded SUSY, where top partners are uncolored.



100 TeV collider can probe these possibilities with exotic Higgs decays/modified Higgs pheno.

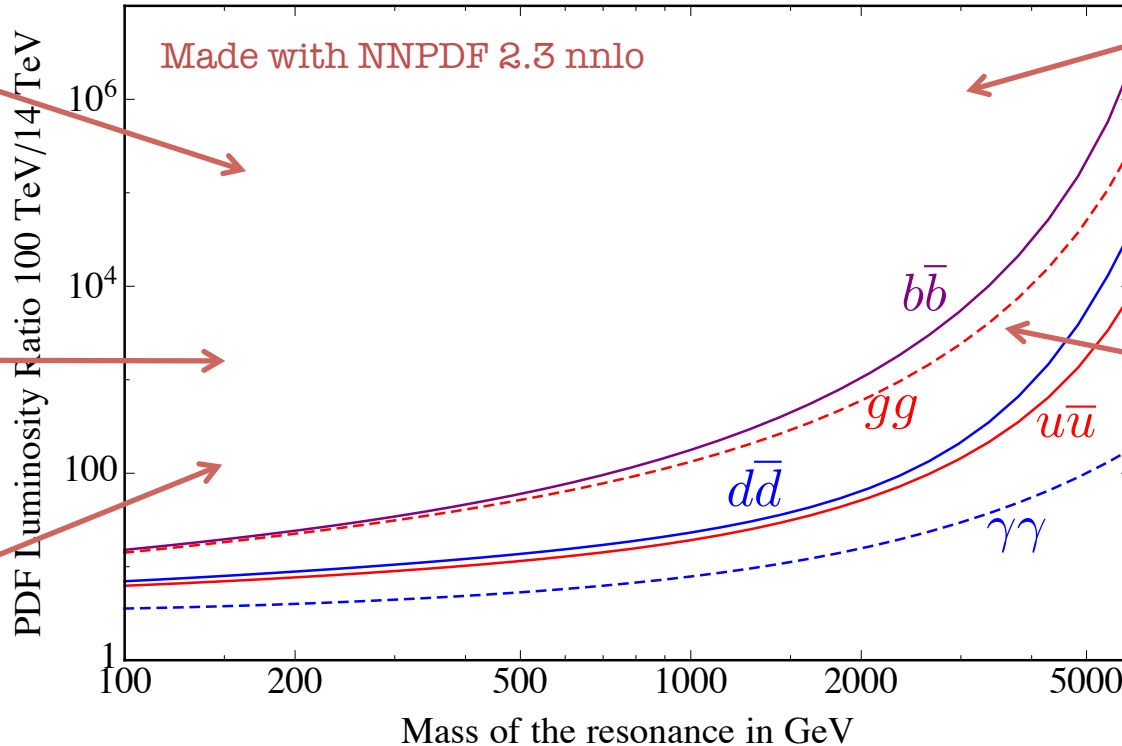
General Comments

It is a very good idea to consider luminosity ratio plots. They tell us about a vast physics program if e.g. a heavy resonance is discovered...

High precision in dominant production modes.

Differential distributions.

Rare/ associated production modes

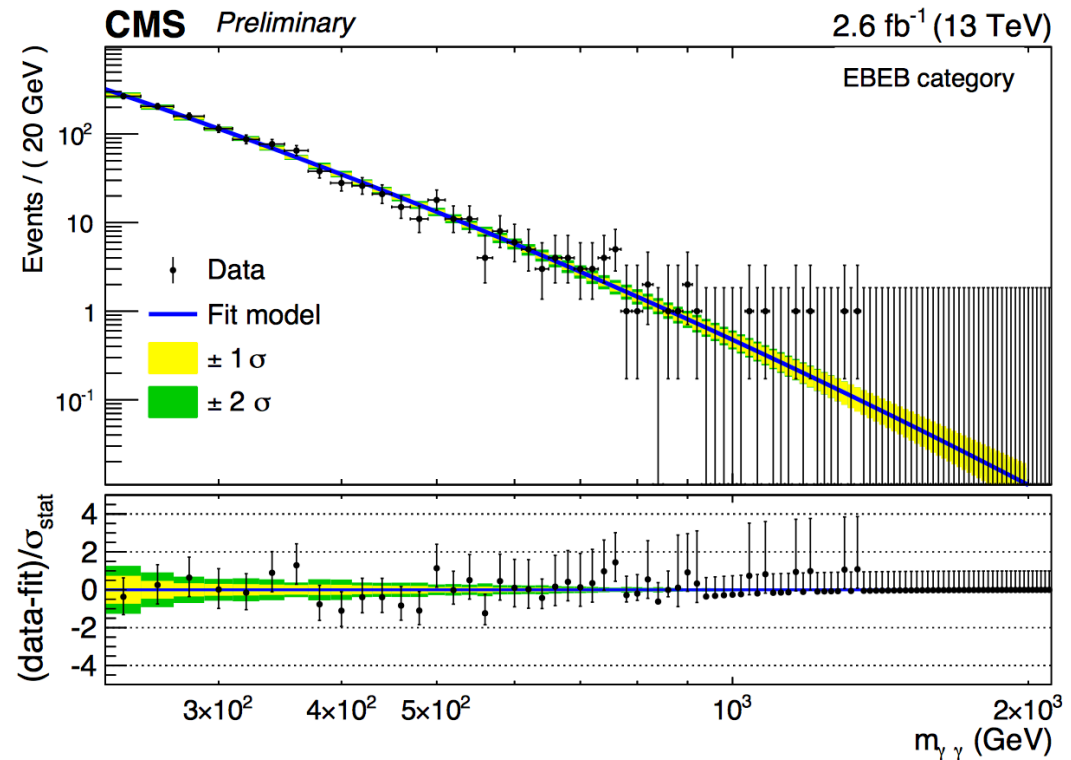
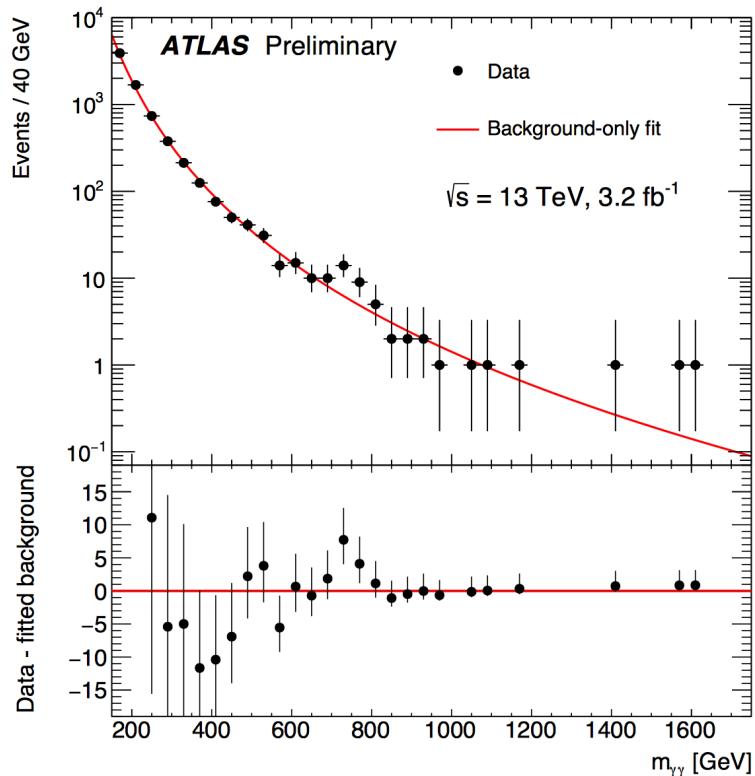


Rare/ exotic decays.

For exotic signatures can take full advantage of cross section if background is small. E.g. displaced vertices.

750 GeV DiPhoton Resonance

ATLAS and CMS both see an excess at 750 GeV in the diphoton spectrum:



First thoughts on what we could do with a 100 TeV collider...

750 GeV DiPhoton Resonance

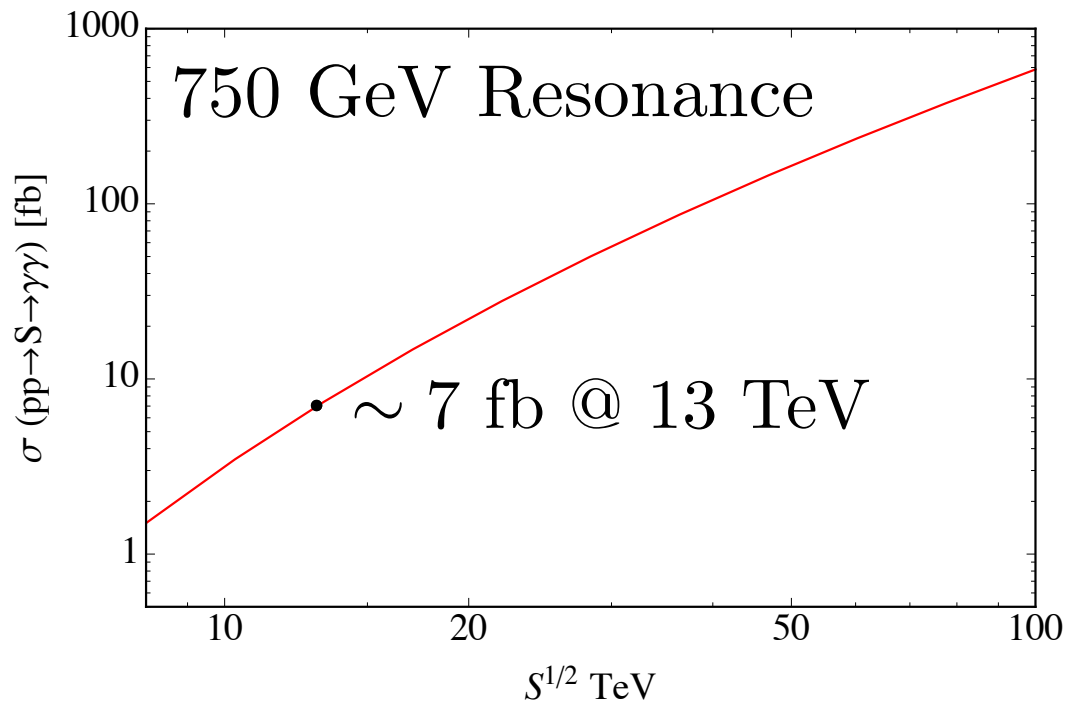
ATLAS and CMS both see an excess at 750 GeV in the diphoton spectrum..

First thoughts on what we could do with a 100 TeV collider...

- If it turns out to be real, then we need to know what we could do with 100 TeV!
- If a statistical fluctuation, nonetheless a good straw man exercise to consider the 100 TeV potential for fully exploring an LHC BSM discovery...

750 GeV DiPhoton Resonance

Going from 13 TeV to 100 TeV huge leap in cross section:



With 3 ab^{-1} at LHC expect

$$N_{\gamma\gamma} \approx 21 \times 10^3$$

however, with the same integrated luminosity at 100 TeV would have

$$N_{\gamma\gamma} \approx 1.8 \times 10^6$$

This would allow for very detailed study indeed!

- Absolute precision
- Distributions
- Rare decays
- Rare production modes

Conclusions

Continued experimental investigation is required to answer many questions in theoretical physics, including

- Electroweak naturalness (and beyond)
- Dark Matter
- ... and much else

A 100 TeV proton collider would explore new territories beyond the Standard Model and, in the event of a discovery at the LHC, could offer precision measurements of LHC discoveries as well as further exploration to uncover underlying structures.